

Measurement of transverse single spin asymmetry at forward rapidity by the STAR experiment in p+p collisions at $\sqrt{s} = 200$ and 500 GeV

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Light Cone 2021: Physics of Hadrons on the Light Front

Outline

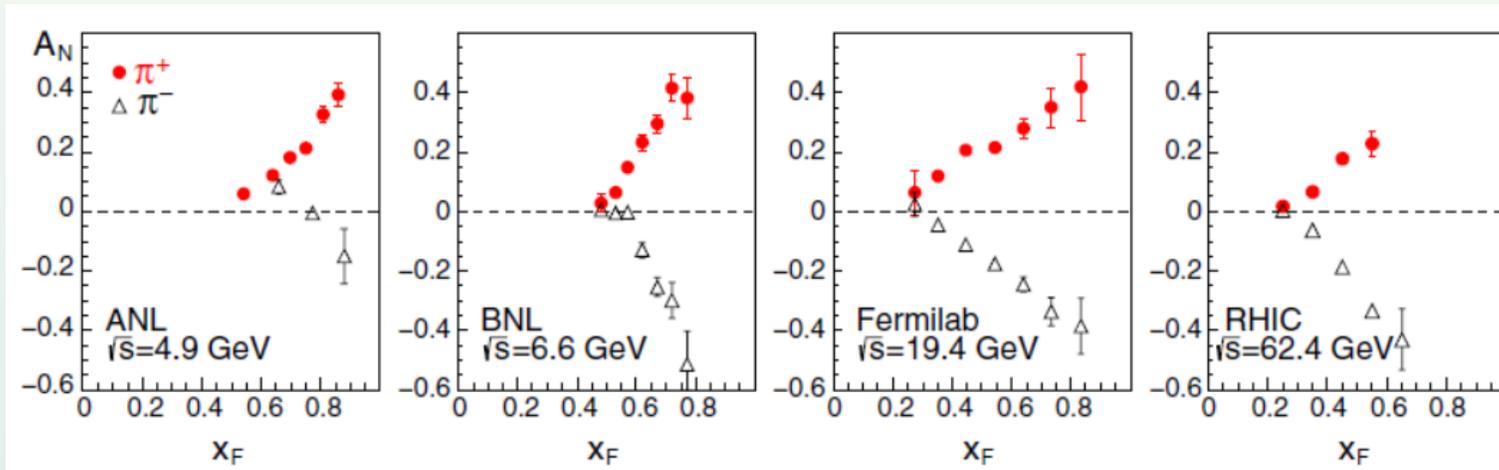
- Motivation
- Experiment setup
- Analysis
- Result and discussion
- Summary

This work has been Published in Phys. Rev. D 103, 092009 in 2021

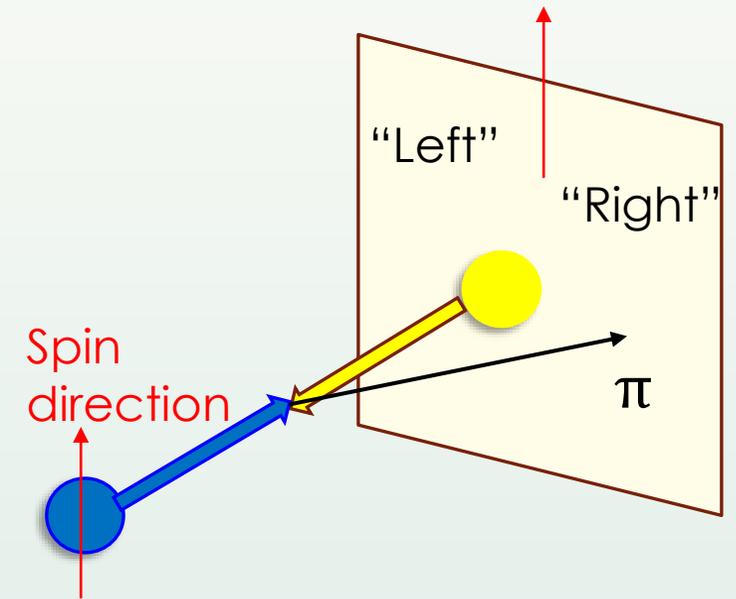
Motivation

- Transverse single spin asymmetry (**TSSA**/ A_N)
- The large forward TSSA was first discovered in 1970s and can not be explained by LO QCD calculation

$$A_N = \frac{d\sigma^\uparrow - d\sigma^\downarrow}{d\sigma^\uparrow + d\sigma^\downarrow}$$



Aidala et al., Rev. Mod. Phys. 85, 655



Motivation

- Great progress has been made over the decades: Transverse momentum dependent PDF (TMD), Collinear twist-3 factorization and etc.

In general, these models share some similarities that the TSSA originates from the **initial and final state effects**

- A decomposition of the contributions to TMD

- **Initial state effect:** asymmetry originates from parton distribution function

$$\hat{f}_{q/p^\dagger}(x, \mathbf{k}_\perp) = f_{q/p}(x, k_\perp) + \frac{1}{2} \Delta^N f_{q/p^\dagger}(x, k_\perp) \mathbf{S} \cdot (\hat{\mathbf{P}} \times \hat{\mathbf{k}}_\perp) \quad \text{Sivers function}$$

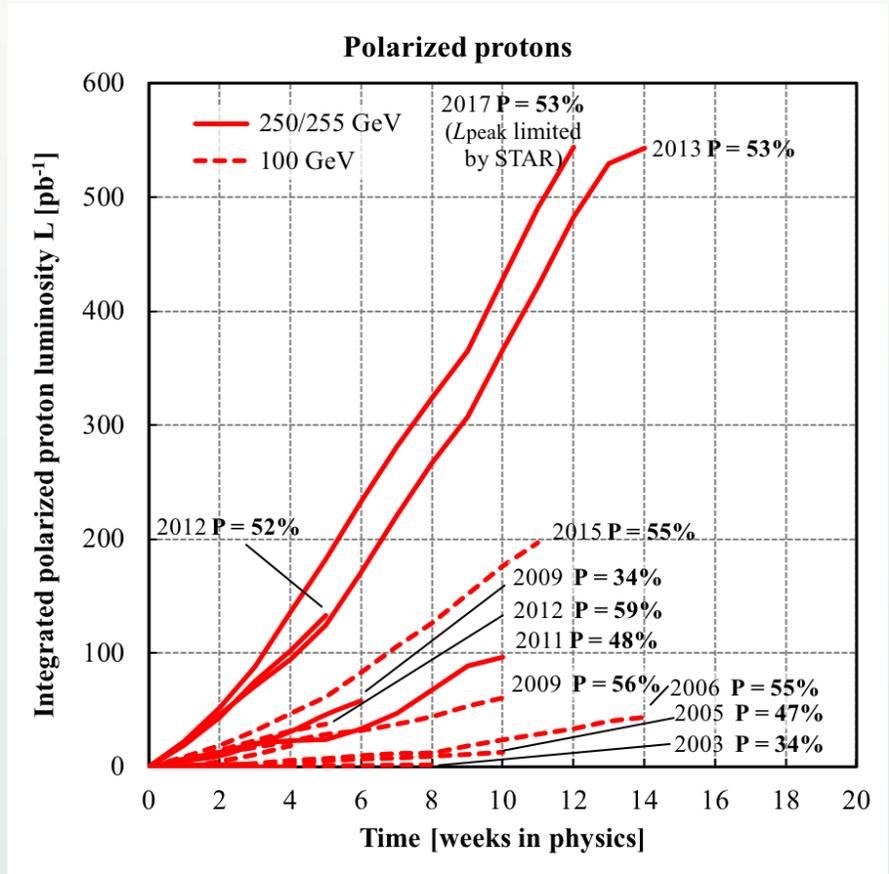
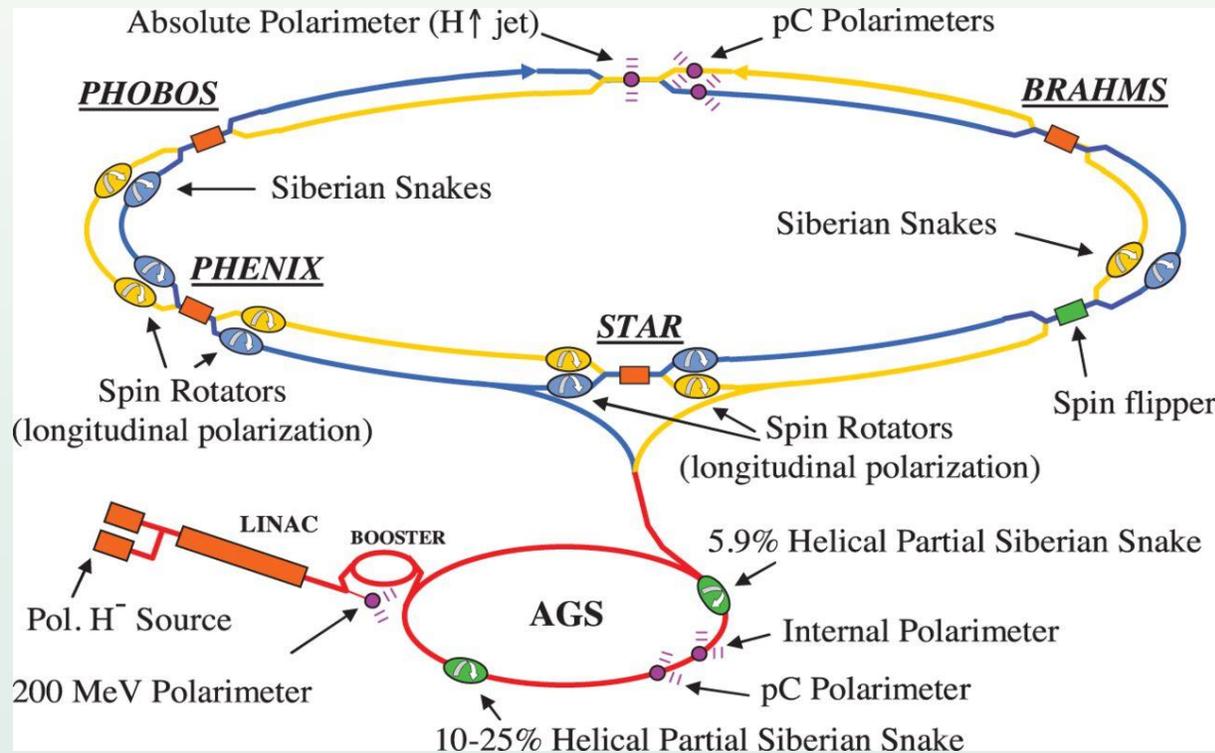
- **Final state effect:** asymmetry originates from fragmentation

Transversity \otimes **Collins function**

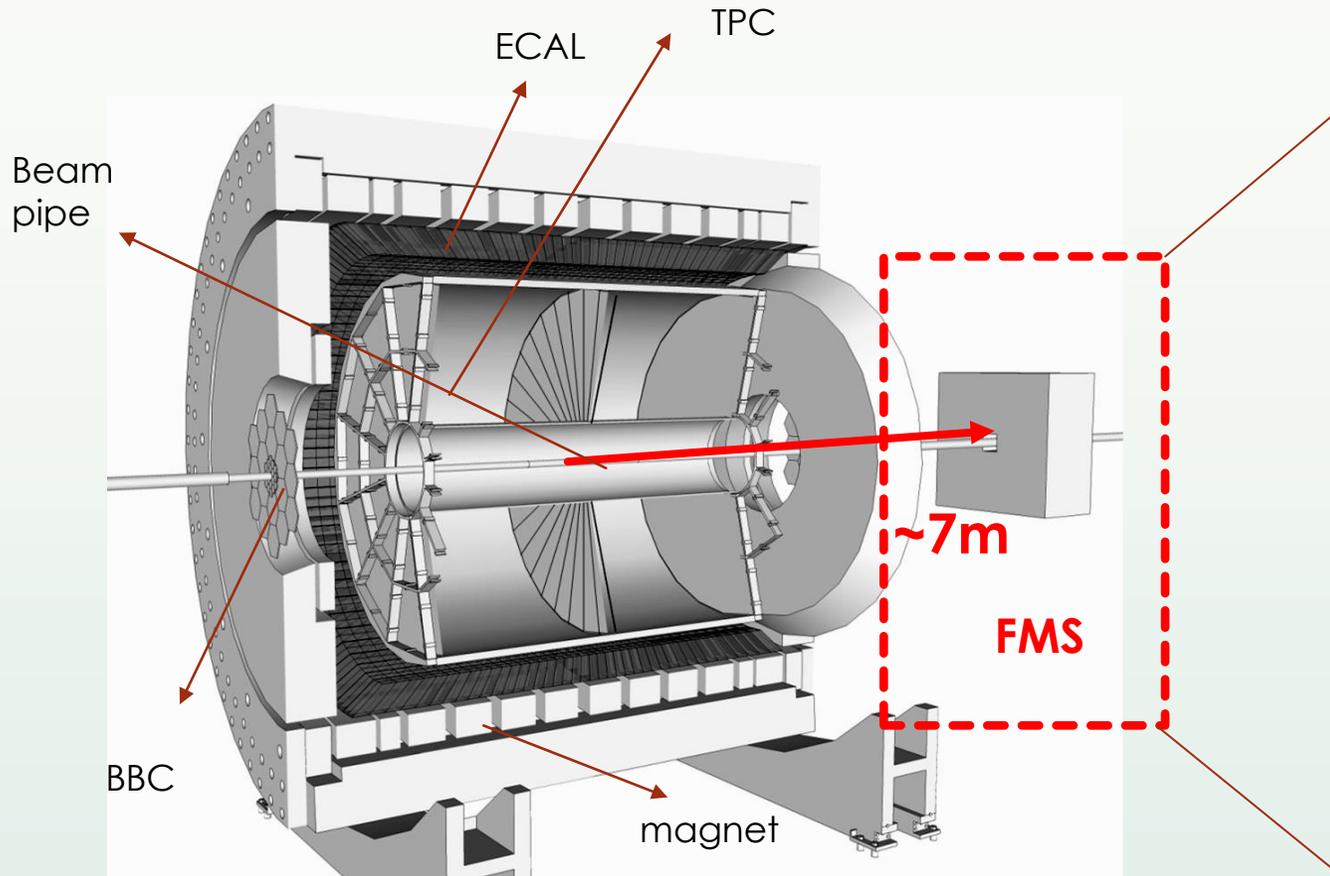
- Experimental data are very important in validating the factorization and constraining the PDFs

Experiment Setup- RHIC & STAR

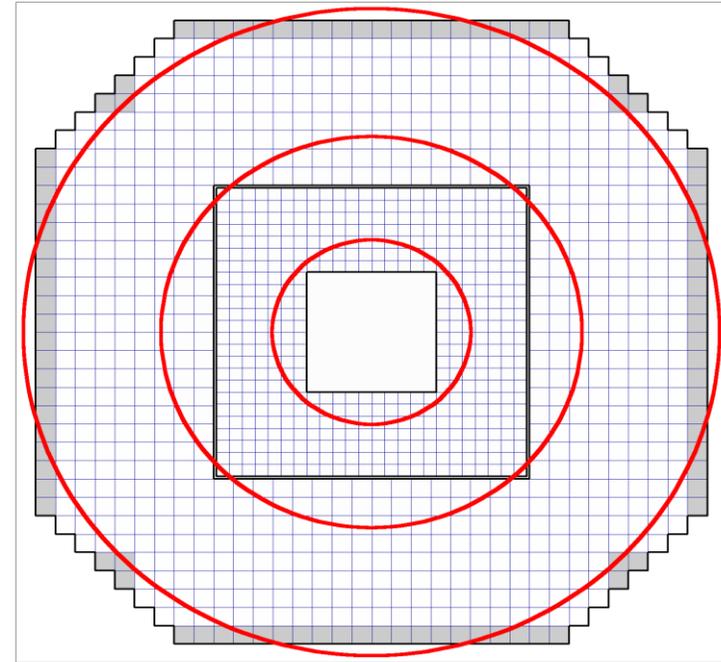
- The **R**elativistic **H**eavy **I**on **C**ollider at BNL provides unique opportunity to study spin physics because it is the world's only polarized proton-proton collider.



Experiment Setup- STAR & FMS



FMS Layout



- ❑ EM-Calorimeter made of 1000+ lead glass cells
- ❑ Large pseudo-rapidity coverage in the forward direction 2.6-4.1
- ❑ Two cell types

Analysis- Dataset

► Dataset:

Transversely polarized proton-proton collisions

Year	Energy	Events
2011	500 GeV	165M
2015	200 GeV	569M

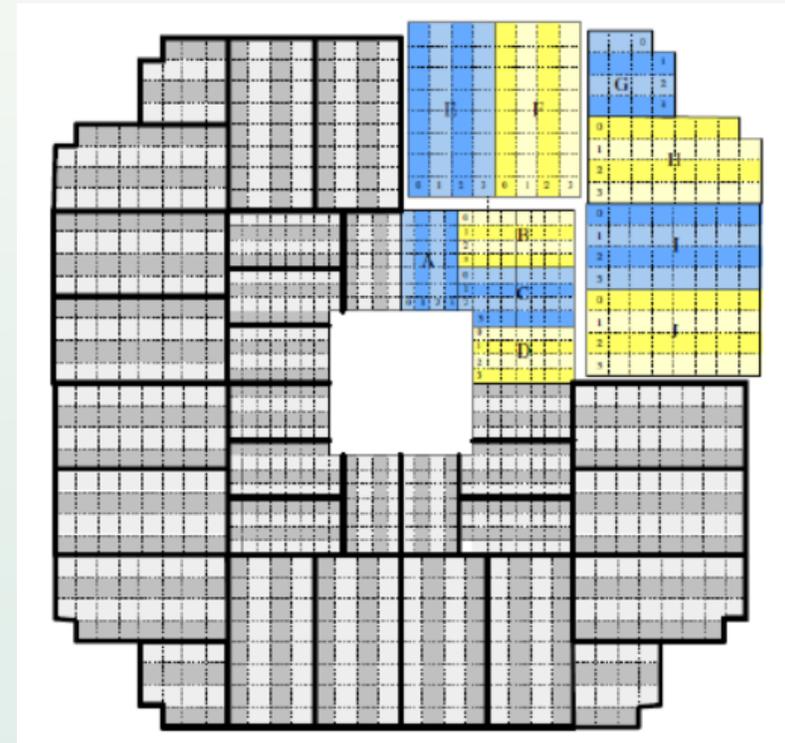
► Beam polarization:

52 / 57% (500 / 200 GeV)

► Trigger:

FMS-Board-sum and FMS-Jet-patch, both based on energy deposition in a defined region of the FMS

Trigger logic



Analysis- Asymmetry calculation

The **luminosity** and **detector efficiency** can be difficult to determine.

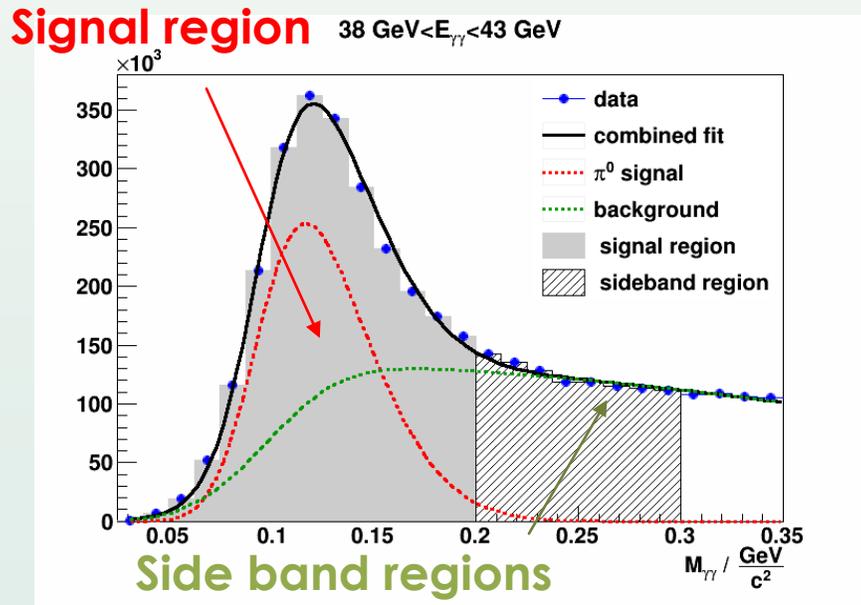
$$N^\uparrow(\phi) = \epsilon \mathcal{L}^\uparrow \sigma^\uparrow$$

$$= \epsilon \mathcal{L}^\uparrow (1 + \text{pol} * A_N \cos \phi) \sigma$$

➔ “Cross-ratio” method help eliminate those factors

$$\text{pol} \cdot A_N^{\text{raw}} \cos \phi = \frac{\sqrt{N^\uparrow(\phi)N^\downarrow(\phi + \pi)} - \sqrt{N^\downarrow(\phi)N^\uparrow(\phi + \pi)}}{\sqrt{N^\uparrow(\phi)N^\downarrow(\phi + \pi)} + \sqrt{N^\downarrow(\phi)N^\uparrow(\phi + \pi)}}$$

Typical di-photon mass spectrum



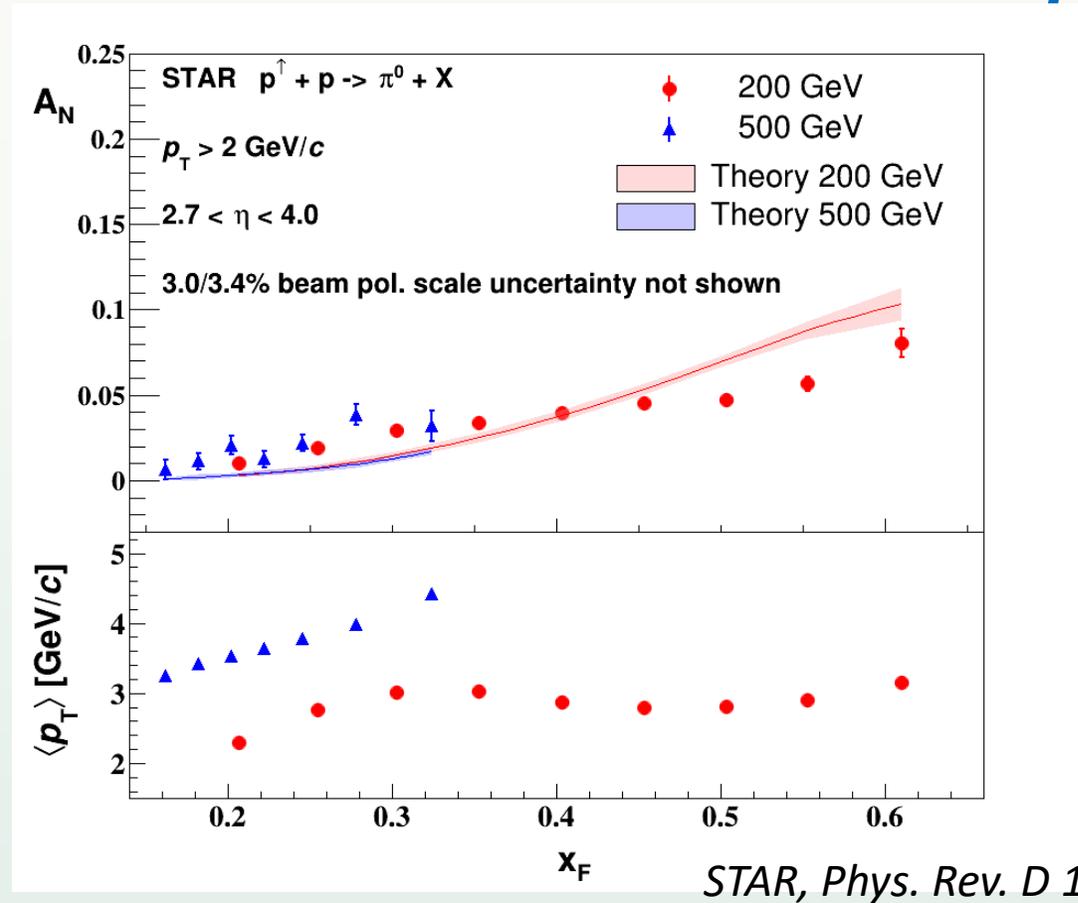
➔ Background subtraction

The fraction comes from the fitting of the mass spectrum
Signal/background shapes are from simulation

$$A_N^{\text{raw}_{sig}} = f_{\text{sig}_{sig}} * A_N^{\pi^0} + (1 - f_{\text{sig}_{sig}}) * A_N^{\text{bkg}}$$

$$A_N^{\text{raw}_{sb}} = f_{\text{sig}_{sb}} * A_N^{\pi^0} + (1 - f_{\text{sig}_{sb}}) * A_N^{\text{bkg}}$$

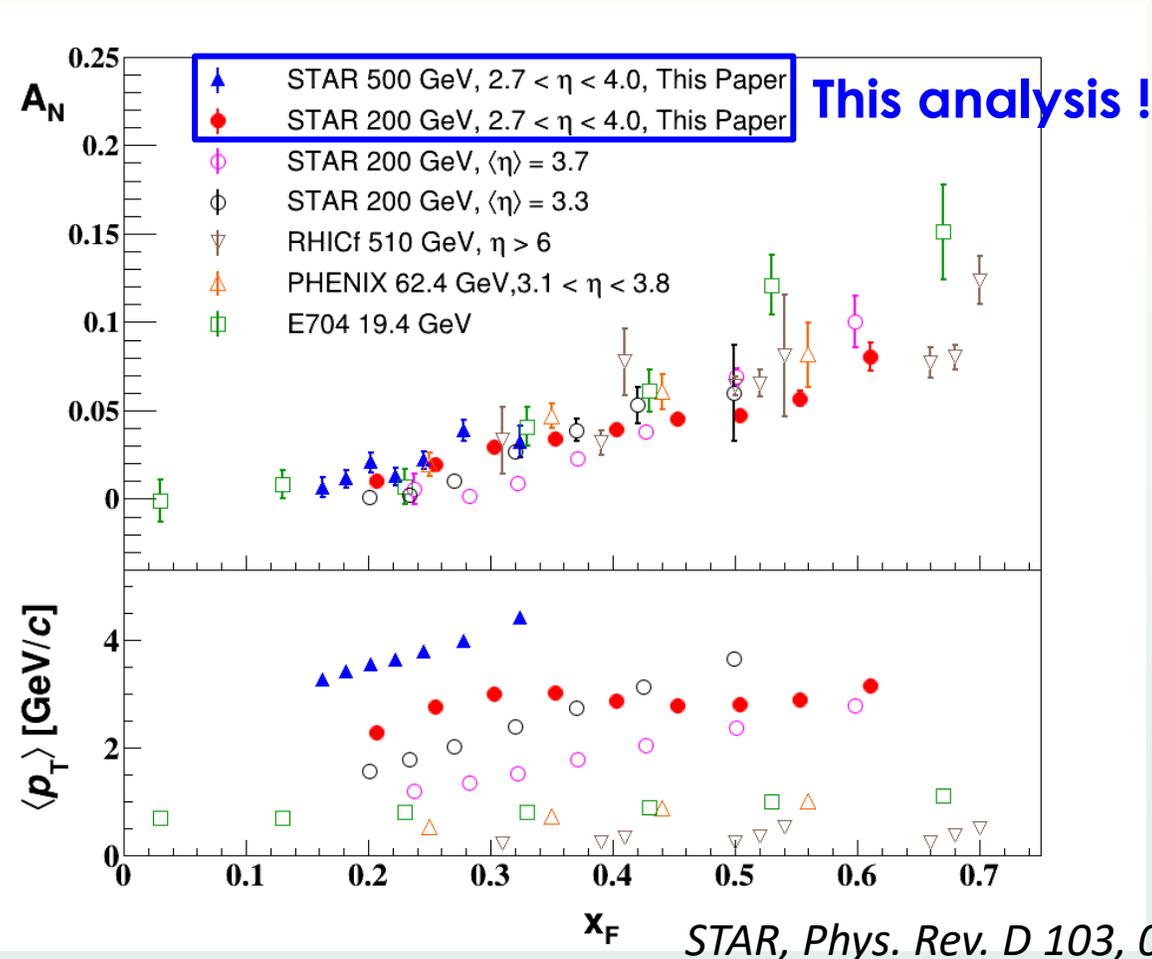
Result- π^0 TSSA vs. x_F



Theory curves:
 J. Cammarota et al.,
 Phys.Rev.D 102, 054002

- ❑ The π^0 TSSA increases with x_F .
- ❑ Consistent between 200 GeV and 500 GeV. Energy dependence is weak.

Comparison to previous measurements



□ Weak collision energy dependence of the π^0 TSSA from 19.4 to 500 GeV

□ Other observables are needed to disentangle the initial and final state effect.

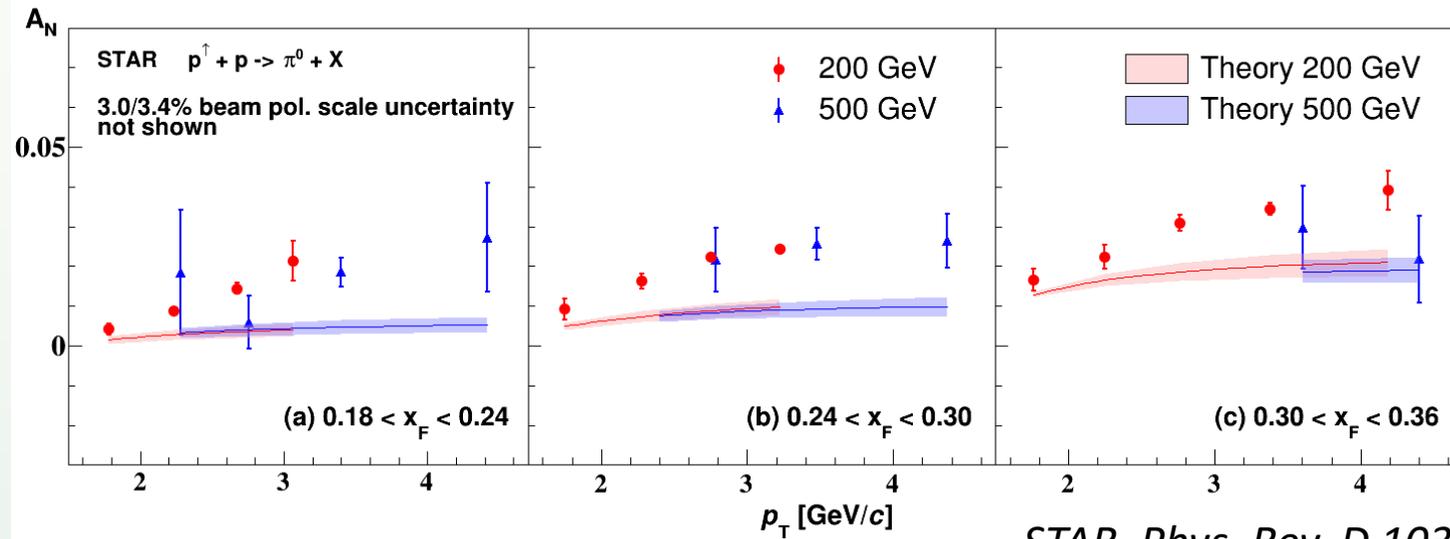
STAR, *Phys. Rev. D* 103, 092009

STAR 2008 data:

STAR, *Phys.Rev.Lett.* 101, 222001

$$x_F = \frac{E_L^{\pi^0}}{E_{beam}}$$

Result- π^0 TSSA vs. p_T



STAR, Phys. Rev. D 103, 092009

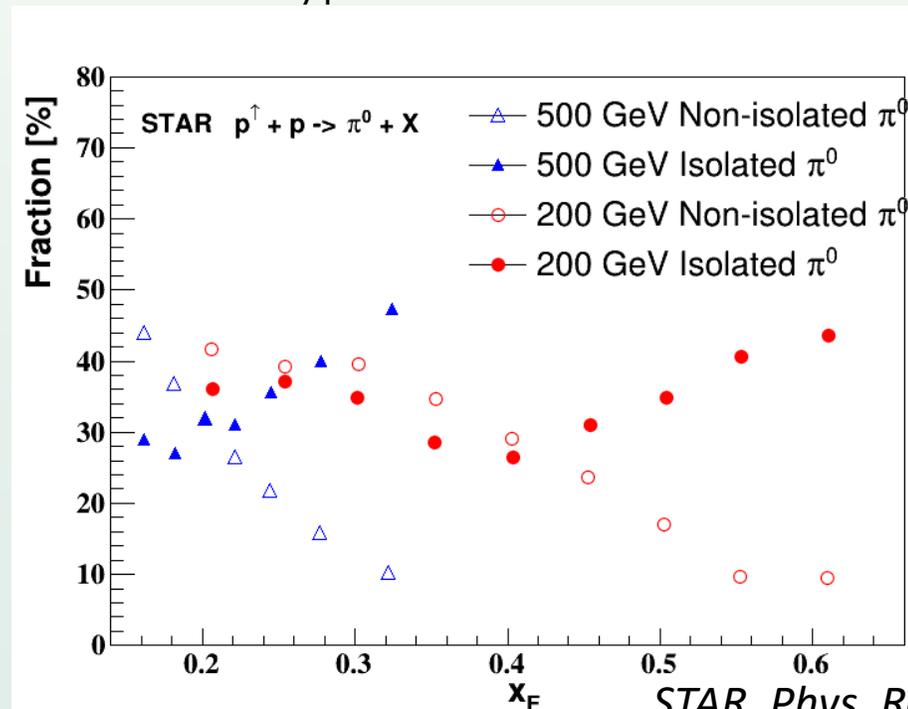
Theory curves:
 J. Cammarota et al.,
 Phys.Rev.D 102, 054002

- ❑ Overlapping x_F region between 200 GeV and 500 GeV results.
- ❑ The 200 GeV data shows a significant increase of TSSA below 3 GeV, which explains the higher TSSA than 2008 STAR data in the previous slide.
- ❑ The 500 GeV data are flat over the p_T range.

Result- isolated π^0 TSSA

- ❑ Motivation: investigate the π^0 event topology (π^0 with no other particle around)
- ❑ Method: in a surrounding area (in η - ϕ space, $R=0.7$), if the π^0 energy fraction $> 98\%$, it is defined as isolated. If the π^0 energy fraction $< 90\%$, it is defined as non-isolated

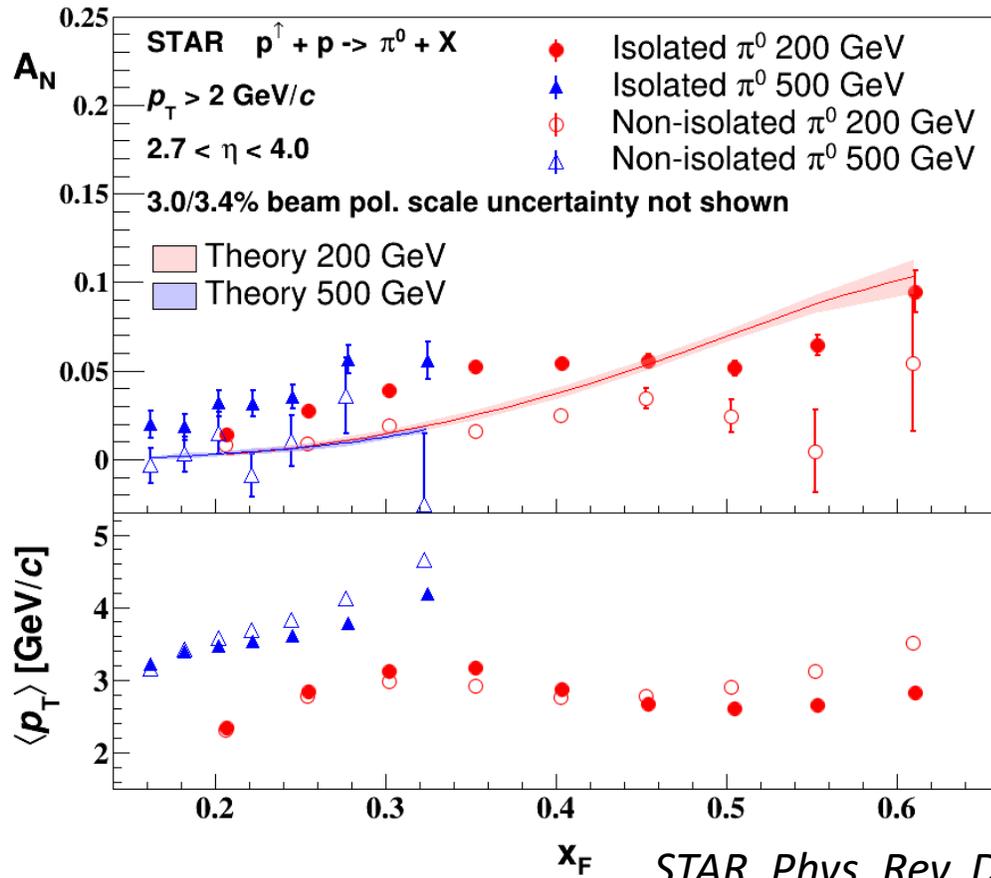
Fractions of different types of π^0 event in the overall sample



STAR, *Phys. Rev. D* 103, 092009

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Result- isolated π^0 TSSA



STAR, *Phys. Rev. D* 103, 092009

Theory curves:

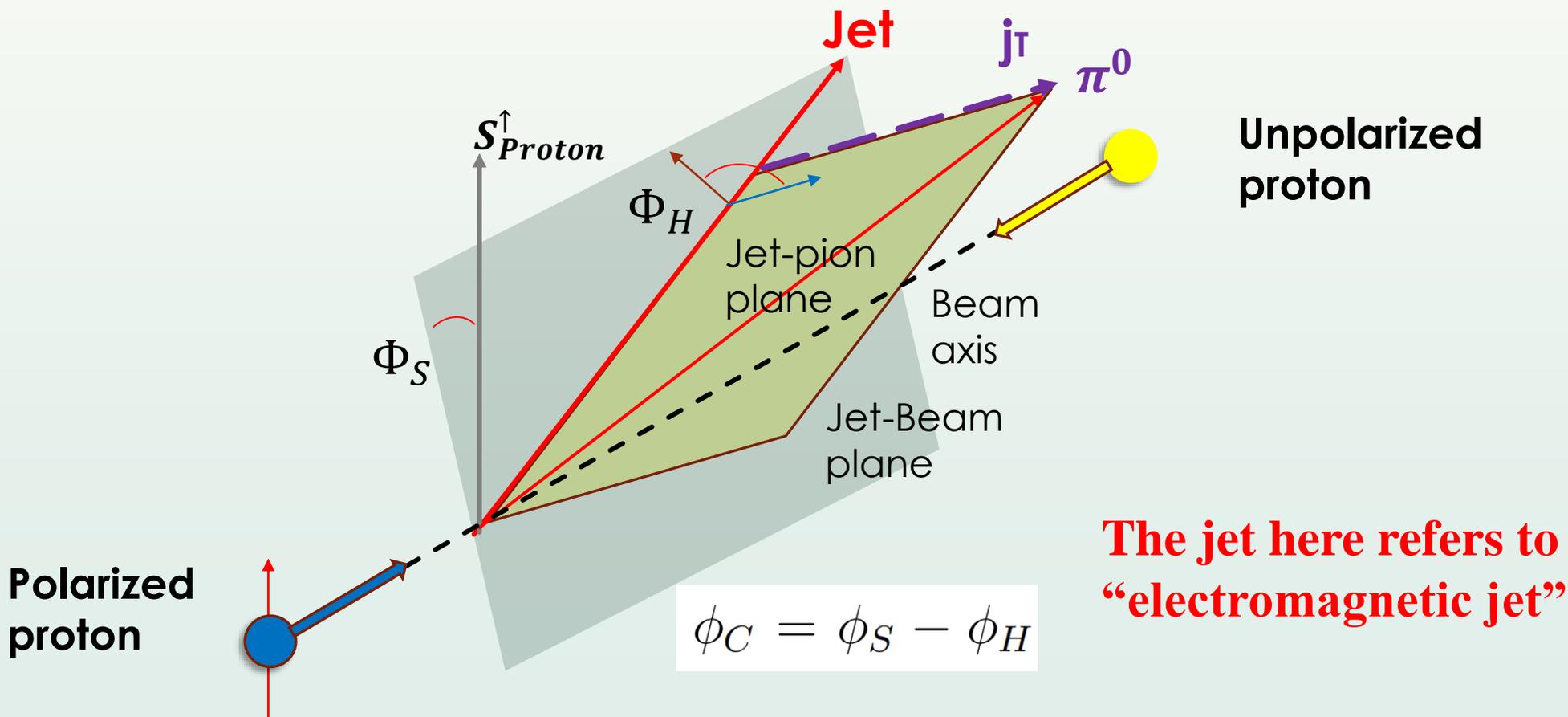
J. Cammarota et al., *Phys.Rev.D* 102, 054002

- The TSSAs of the two types of π^0 are significantly different. Isolated π^0 TSSA dominates.
- The physical origin and mechanism accounting for higher TSSA of isolated π^0 is not known yet – implication of a third origin? Diffractive process is a potential candidate.

More observables

Jet TSSA – sensitive to the initial state effect.

Collins asymmetry – sensitive to the final state effect.



More observables

π^0 /EM-jet TSSA

vs.

Collins asymmetry

$$\begin{aligned} N^\uparrow(\phi) &= \epsilon \mathcal{L}^\uparrow \sigma^\uparrow \\ &= \epsilon \mathcal{L}^\uparrow (1 + \text{pol} * A_N \cos \phi) \sigma \end{aligned}$$

$$\begin{aligned} N^\uparrow(\phi_c) &= \epsilon \mathcal{L}^\uparrow \sigma^\uparrow \\ &= \epsilon \mathcal{L}^\uparrow (1 + \text{pol} * A_{UT} \sin \phi_c) \sigma \end{aligned}$$

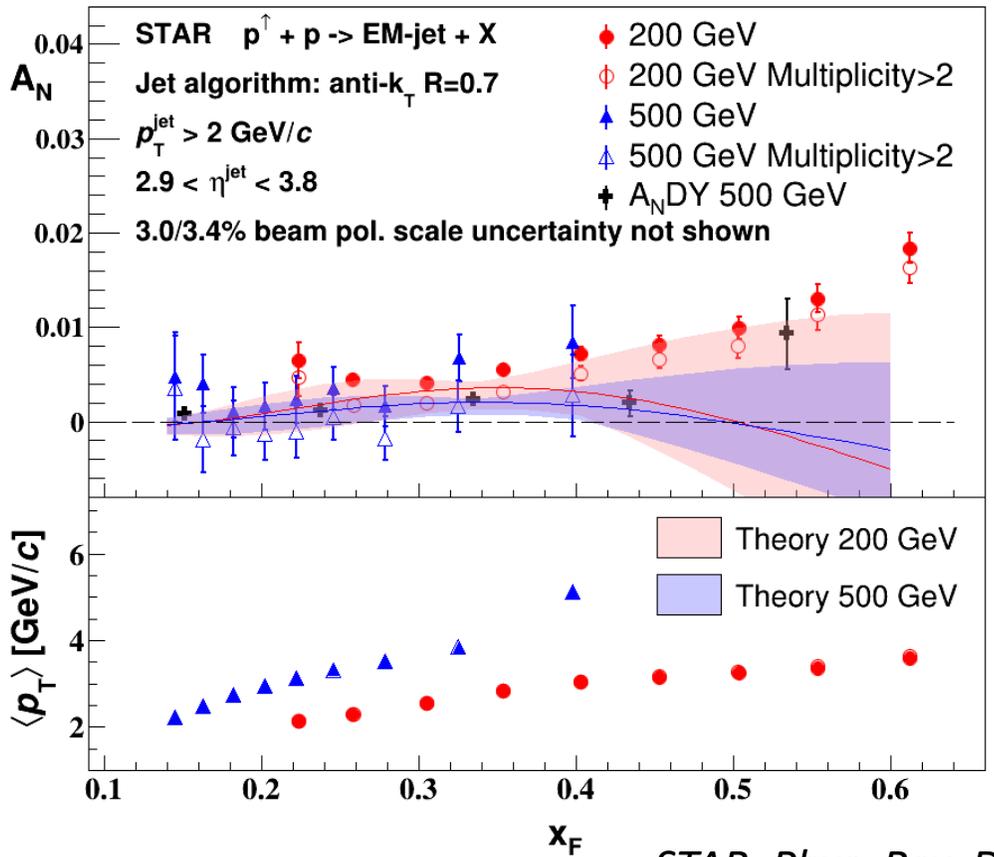
- Azimuthal angle ϕ
- All π^0 candidates
- Background subtraction for π^0

- Collins angle ϕ_c
- Only π^0 within a jet
- No background subtraction

For jet reconstruction and π^0 in a jet :

- anti- k_T $R=0.7$
- $p_T > 2$ GeV
- $\sqrt{(\phi_{\pi^0} - \phi_{jet})^2 + (\eta_{\pi^0} - \eta_{jet})^2} < 0.04$

Result- jet TSSA



STAR, Phys. Rev. D 103, 092009

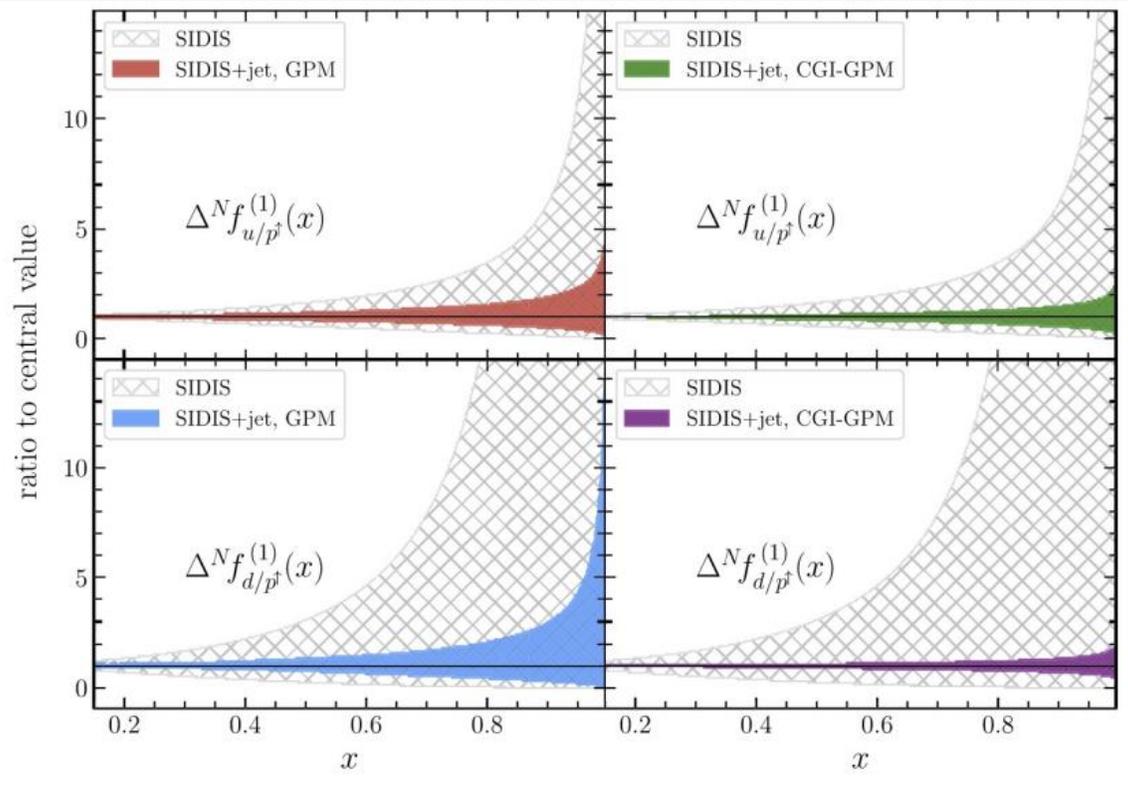
Theory curves:

L. Gamberg et al., Phys.Rev.Lett. 110,232301

- The jet TSSA is a few times smaller than the π^0 TSSA in the same x_F bin.
- Jets with minimum photon multiplicity requirement have significantly smaller TSSA.
- The A_N DY result shows the TSSA of full jets, and is consistent with the result of the EM-jets having at least 3 photons.

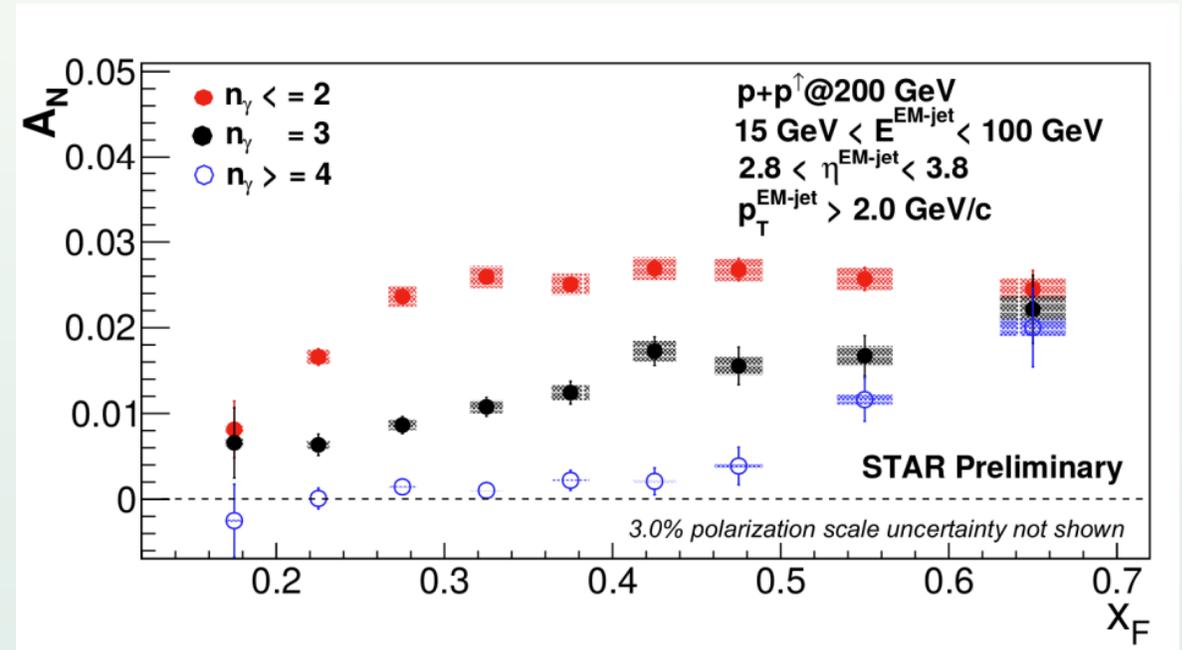
Result- jet TSSA

- Impact: greatly reduce the uncertainty of the u/d quark Sivers function at large x

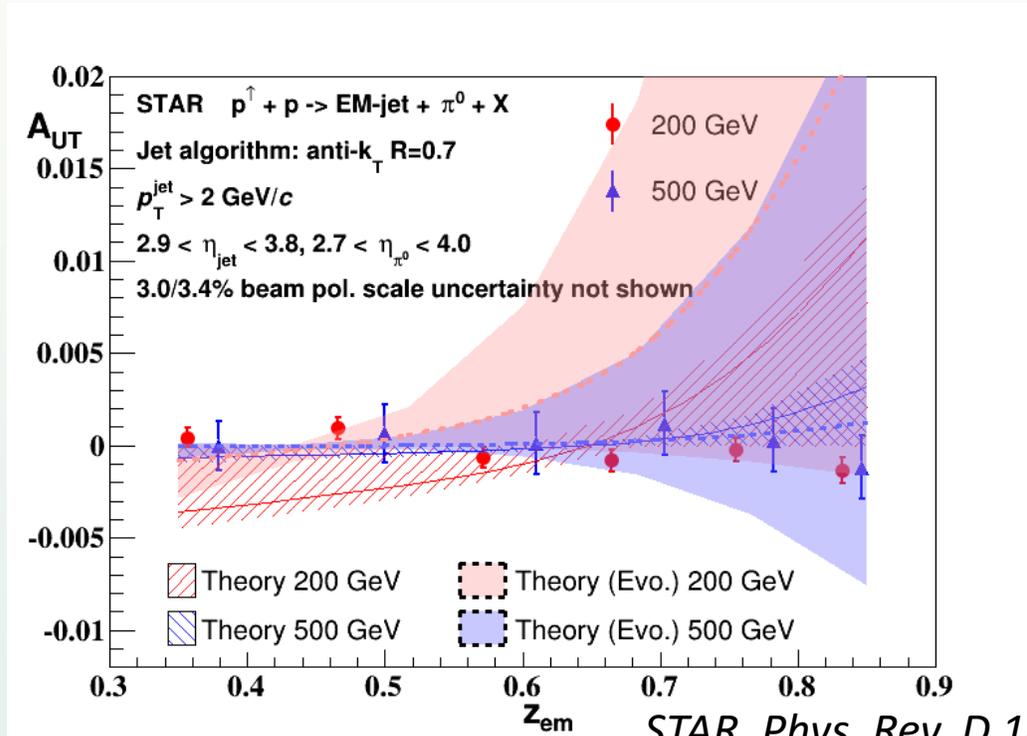


M. Boglione et al., Phys.Lett.B 815, 136135

- Follow up work at STAR: EM-jet TSSA decreases with increasing photon multiplicity (more “jet-like”)



Result- Collins asymmetry for π^0 in a jet



□ The Collins asymmetries are very small at both energies, which reflects the cancellation of the Collins effect of the u/d quark

$$z_{em} = \frac{E_\pi}{E_{jet}}$$

STAR, *Phys. Rev. D* 103, 092009

Theory curves:

Z. Kang et al., *Phys.Lett.B* 774,635

The whole landscape

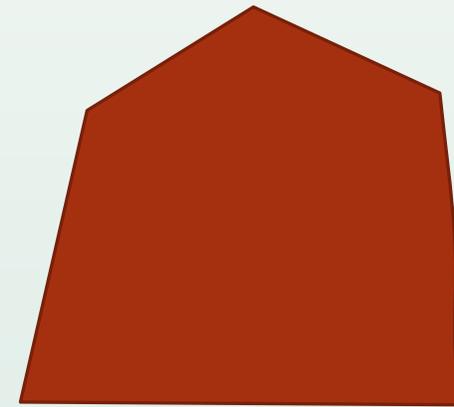
The π^0 TSSAs have multiple sources. Interesting things happen when we break it down:



$$\pi^0 \text{ TSSA} = \text{Initial state effect} + \text{Final state effect} + \text{Unclear source}$$



Isolated π^0 TSSA



Jet TSSA



Initial state effect

Collins asymmetry



Final state effect

Unclear source
(Diffractive process?)

Summary

- ❑ We measured the π^0 /jet TSSA and Collins asymmetry using the FMS in STAR 200 and 500 GeV p-p data
- ❑ We investigated the π^0 event topology. The mechanism for the higher TSSAs of the isolated π^0 remains unclear. It offers new perspectives to the origin of TSSA
- ❑ We measured the jet TSSAs and Collins asymmetry to separate contributions from initial and final state effects, both of which are small
- ❑ These measurements together provide important inputs for further investigation for TSSA
- ❑ This work has been Published in Phys. Rev. D 103, 092009 in 2021